

Multipole Hair of Higher Dimensional Schwarzschild Black Holes

Based on arXiv:1907.07622 [gr-qc] (Journal of Mathematical Physics)

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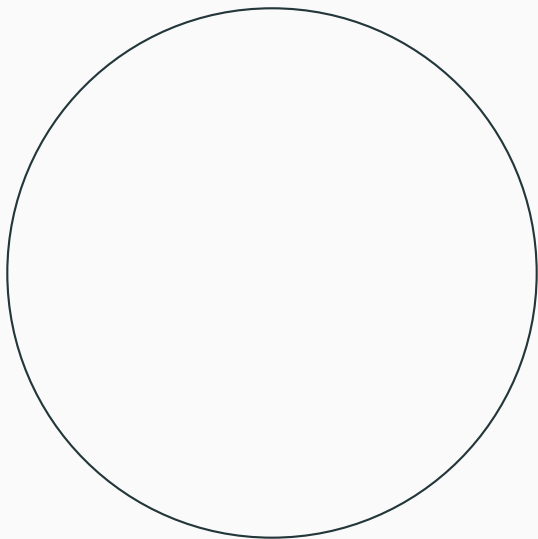
msfox@g.hmc.edu

APS Far West Section, Stanford University

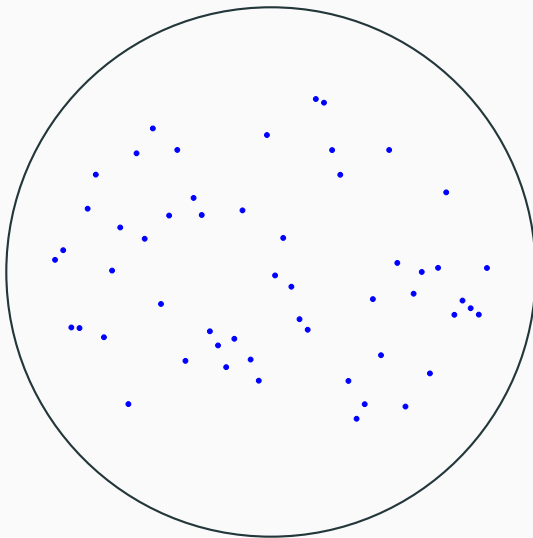
2 November 2019



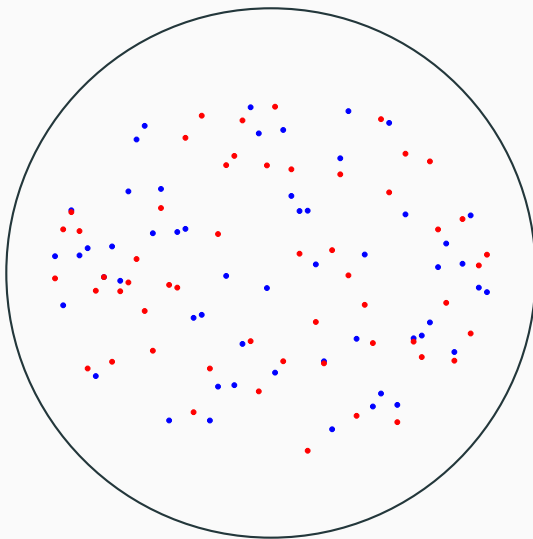
To Be or Not to Be (a Black Hole)



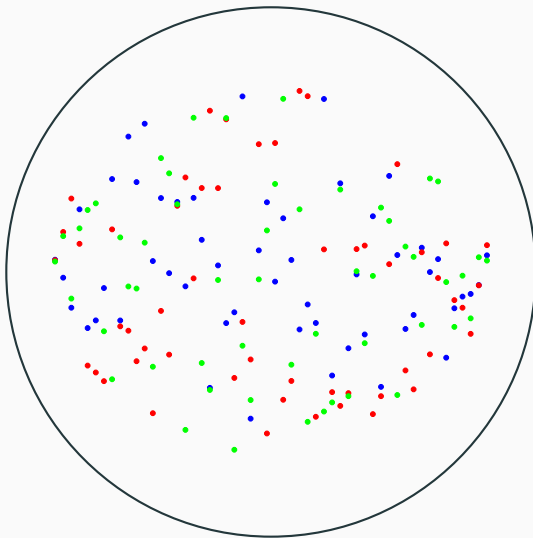
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No hair!

Black Hole Topology in $d \geq 4$ Dimensional Spacetime

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Of course,

$$\partial\mathcal{B}_4 \cong S^2 \not\Rightarrow \partial\mathcal{B}_d \cong S^{d-2}$$

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Theorem (Galloway *et al.*, 2006). Assuming dominant energy condition,

$\partial\mathcal{B}_d$ admits metric of positive scalar curvature

G. J. Galloway and R. Schoen, *Commun. Math. Phys.* 266, 571 (2006)

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Example (Black Saturn). 5D black hole $\mathcal{B}_5$ with horizon topology

$$\partial\mathcal{B}_5 \cong S^2 \times S^1$$

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So, higher dimensional black holes not necessarily unique.

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Theorem (Hwang, 1998). All topologically-spherical, static, asymptotically-flat, and non-degenerate vacuum solutions to Einstein equations have ST geometry, g_{ST}

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Theorem (Gibbons *et al.*, 2002). All topologically-spherical, static, asymptotically-flat, and non-degenerate electrovac solutions to Einstein-Maxwell equations have RNT geometry, g_{RNT}

Dropping Electric Charge into ST Black Hole: $g_{\text{ST}} \rightarrow g_{\text{RNT}}$?

Question: How to get RNT black hole from ST black hole?

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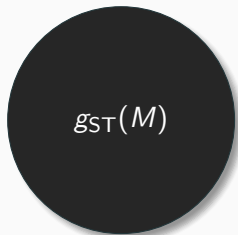
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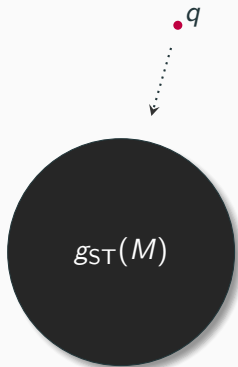
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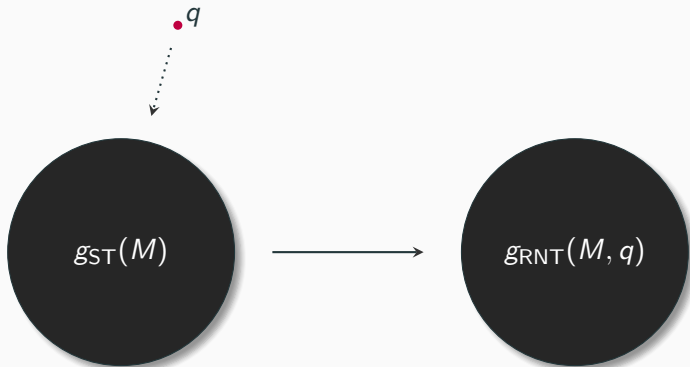
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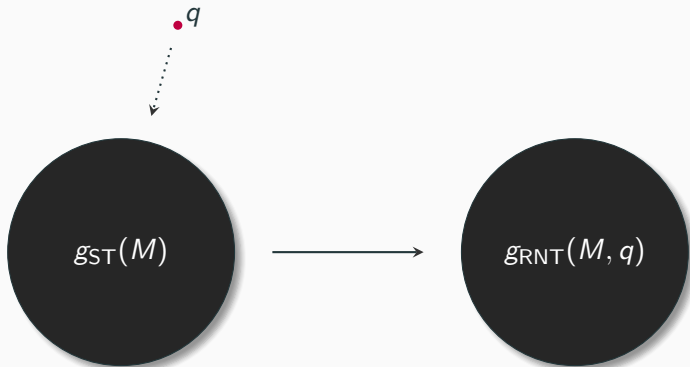
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$$\Phi_{\text{RNT}}(r) \neq \Phi(r) \implies \boxed{g_{\text{ST}}(M) \rightarrow g(M, q) \neq g_{\text{RNT}}(M, q)}$$

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Thus,

final state is not topologically spherical

Acknowledgements

- Brian Shuve (Harvey Mudd College)
- Jason Gallicchio (Harvey Mudd College)
- Thomas Moore (Pomona College)